$^{26}\text{Si} + \alpha$ resonant scattering measurement to study $^{26}\text{Si}(\alpha, p)^{29}\text{P}$ reaction rate

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An X-ray burst can be characterized by a sudden and intense release of X-ray radiation from a compact stellar object such as a neutron star. A total energy release of approximately 10^{39-40} ergs can be achieved per burst in just a few seconds. It is believed that the proton-rich nuclei up to the Sn-Sb-Te region can be synthesized during the burst. To better understand the X-ray bursts, studying the ${}^{26}\text{Si}(\alpha, p){}^{29}\text{P}$ reaction is essential since ${}^{26}Si$ is considered to be one of the waiting points in the nucleosynthesis.

A sensitivity study, which identifies the important nuclear reaction rates that affect the X-ray light curves or ash composition, suggests that the ${}^{26}Si(\alpha, p){}^{29}P$ is one of the impactful reactions to the light curve of the X-ray burst.¹⁾ Despite its importance, the study of ${}^{26}\mathrm{Si}(\alpha,p){}^{29}\mathrm{P}$ is less understood experimentally. Thus we have performed a ${}^{26}Si + \alpha$ experiment to measure the ${}^{26}Si(\alpha, p){}^{29}P$ reaction directly and the resonant scattering to investigate resonances in ³⁰S, which can be populated in the ${}^{26}Si(\alpha, p){}^{29}P$ reaction as intermediate states. The result on the ${}^{26}\text{Si} + \alpha$ resonant scattering is described in this report.

The ²⁶Si + α resonant scattering was measured at the Center for Nuclear Study Radioactive Ion Beam Separator $(CRIB)^{2}$ of the University of Tokyo. The radioactive ²⁶Si beam was produced at E =2.14 MeV/nucleon through the ${}^{3}\text{He}({}^{24}\text{Mg}, n){}^{26}\text{Si}$ reaction by impinging ²⁴Mg at E = 7.56 MeV/nucleon on a cryogenic ³He gas target.³) The ²⁶Si beam was separated and purified by combining the magnetic analysis and velocity selection with a double achromatic system and a Wien filter. Two PPACs were located at the upstream of the reaction target for the event-by-event monitoring of beam position and time-of-flight. The typical ²⁶Si beam intensity was 2.8×10^4 pps, and the beam purity was $\sim 16\%$. The ²⁶Si beam was impinged on the ⁴He gas target with a pressure of 250 Torr. The reaction target was kept at room temperature. The thick target method is adopted for the experiment to scan a wide excitation energy range in 30 S.

The light charged particles were measured by silicon detector telescopes. Using four layers of silicon

detectors ($\Delta E1$, $\Delta E2$, E1, and E2 layer), each species of charged particles could be easily identified. The $\Delta E1$ and $\Delta E2$ detectors are segmented into 16 strips providing the horizontal and vertical position information, respectively. The E1 and E2 detectors are pad type silicon detectors. The scattering angle was obtained based on the position information. To obtain the excitation function of ${}^{26}\text{Si}(\alpha, \alpha){}^{26}\text{Si}$ reaction, the α energies were converted to the center-of-mass energy by considering the kinematics of the reaction and the energy loss of particles in the gas target. Figure 1 shows the excitation function obtained at $\theta_{C.M.}$ $= 174^{\circ}$. Fitting the experimental excitation function with the theoretical R-matrix calculations, we will extract resonance parameters of levels in the ³⁰S, such as excitation energy, spin, parity, and α partial width, to constrain the ${}^{26}\text{Si}(\alpha, p){}^{29}\text{P}$ reaction rate.



Fig. 1. Excitation function of ${}^{26}\text{Si} + \alpha$ elastic scattering at $\theta_{c.m.} = 174^{\circ}$

References

- 1) R. H. Cyburt et al., Astrophys. J. 55, 830 (2016).
- 2) Y. Yanagisawa et al., Nucl. Instrum. Methods Phys. Res. A, 539, 74 (2005).
- 3) H. Yamaguchi et al., Nucl. Instrum. Methods Phys. Res. A 589, 150 (2008).

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