

Calibration and simulation of LaBr₃:Ce scintillator for analyzing ⁸Li(α , n)¹¹B-reaction data

Y. Mizoi,^{*1,*2,*3} H. Baba,^{*3} A. Bracco,^{*4,*5} F. Camera,^{*4,*5} S. M. Cha,^{*6} K. Y. Chae,^{*6} S. Cherubini,^{*3,*7,*8} H. S. Choi,^{*9} N. N. Duy,^{*6} T. Fukuda,^{*3,*10} S. Hayakawa,^{*2} Y. Hirayama,^{*11} N. Imai,^{*2} H. Ishiyama,^{*3} A. Kim,^{*12} D. H. Kim,^{*6} N. Kitamura,^{*2} S. Kubono,^{*2,*3} M. S. Kwag,^{*6} S. Michimasa,^{*2} M. Mihara,^{*3,*13} H. Miyatake,^{*11} S. Ota,^{*2} R. G. Pizzone,^{*7} H. Shimizu,^{*2,*3} N. K. Uyen,^{*6} R. Wakabayashi,^{*3,*13} Y. X. Watanabe,^{*11} O. Wieland,^{*4} H. Yamaguchi,^{*2} L. Yang,^{*2} N. Zhang,^{*2,*14} and Z. C. Zhang^{*14}

We performed an experiment¹⁾ using LaBr₃:Ce scintillators²⁾ at CRIB in 2018 for investigating controversial large differences among the previously reported³⁻⁹⁾ cross-section data of the ⁸Li(α , n)¹¹B reaction. In order to resolve the discrepancies among them, it is essential to obtain an accurate γ -ray spectrum of ¹¹B; therefore, accurate energy calibration and precise efficiency estimation play a crucial role. We performed off-line measurements for energy calibration using radiation sources, as listed in Table 1. Figure 1 presents the calibration plot, in which the horizontal and vertical axes correspond to the time-to-digital converter (TDC) channel reduced by a charge-to-time converter (QTC) and the expected γ -ray energy, respectively; the plot shows sufficiently good linearity. On-line measurements using ¹⁶N beams were also performed at CRIB. In the studied reaction, ¹⁶N decays to ¹⁶O*, which emits γ -rays with $E_\gamma = 6130$ keV. GEANT4 simulations were also performed for each type of γ -rays. One example is shown in Fig. 2, which is the preliminary result of the γ -ray energy spectrum of ¹⁶N($\beta\gamma$). The histogram with error bars and filled histogram are measured and simulated spectra, respectively. The measured peaks of full energy,

single escape and double escape are consistent with the simulated ones. By applying the results of our precise energy calibration, we can proceed to the data analysis for solving the discrepancies among the previous results.

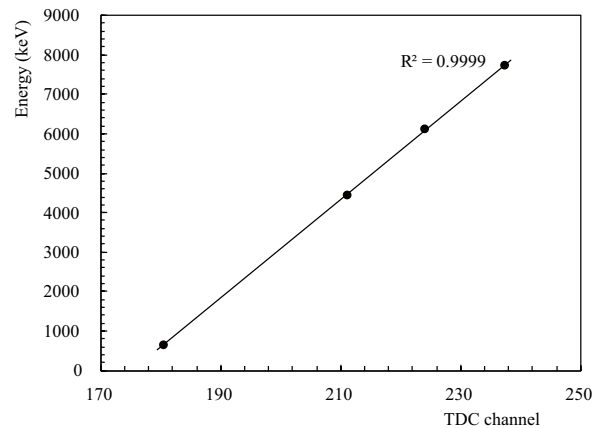


Fig. 1. Calibration plot obtained by off-line measurements.

Table 1. Gamma-ray sources used for off-line calibration and their energies.

Source	Energy (keV)
¹³⁷ Cs	662
Am/Be(¹² C*)	4440
Cm/C(¹⁶ O*)	6130
Am/Be+Al	7724

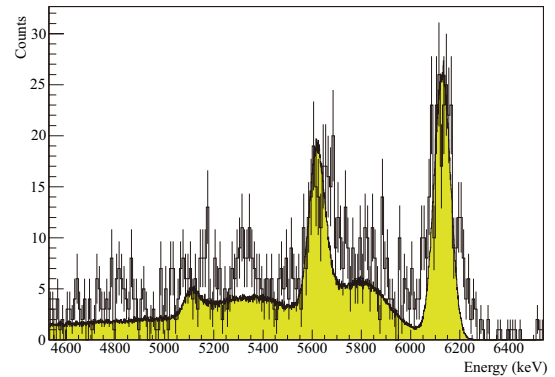


Fig. 2. Energy spectrum of γ -rays from ¹⁶N($\beta\gamma$).

^{*1} Center for Physics and Mathematics, Institute for Liberal Arts and Sciences, Osaka Electro-Communication University
^{*2} Center for Nuclear Study, the University of Tokyo
^{*3} RIKEN Nishina Center
^{*4} INFN - Sezione di Milano
^{*5} Dipartimento di Fisica dell'Università degli Studi di Milano
^{*6} Department of Physics, Sungkyunkwan University
^{*7} INFN - Laboratori Nazionali del Sud (LNS)
^{*8} Dipartimento di Fisica ed Astronomia, Università di Catania
^{*9} Department of Physics and Astronomy, Seoul National University
^{*10} RCNP, Osaka University
^{*11} Wako Nuclear Science Center (WNSC/IPNS), KEK
^{*12} Department of Physics, Ewha Womans University
^{*13} Department of Physics, Osaka University
^{*14} IMP, Chinese Academy of Science (CAS)

References

- 1) Y. Mizoi *et al.*, RIKEN Accel. Prog. Rep. **52**, 58 (2019).
- 2) A. Gaiz *et al.*, Nucl. Instrum. Methods Phys. Res. A **729**, 910 (2013).
- 3) R. N. Boyd *et al.*, Phys. Rev. Lett. **68**, 1283 (1992).
- 4) X. Gu *et al.*, Phys. Lett. B **343**, 31 (1995).
- 5) Y. Mizoi *et al.*, Phys. Rev. C **62**, 065801 (2000).
- 6) S. Cherubini *et al.*, Eur. Phys. J. **20**, 355 (2004).
- 7) H. Ishiyama *et al.*, Phys. Lett. B **640**, 82 (2006).
- 8) M. La Cognata *et al.*, Phys. Lett. B **664**, 157 (2008).
- 9) S. K. Das *et al.*, Phys. Rev. C **95**, 055805 (2017).